

# Mass Spectrum and Dark Matter in the CSE<sub>6</sub>SSM

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New Directions in Subatomic Physics

# The MSSM

- ▶ MSSM at tree-level:  $m_{h_1}^2 \leq M_Z^2 \cos^2 2\beta \lesssim (91 \text{ GeV})^2$
- ▶  $m_{h_1} \approx 125.09 \text{ GeV} \Rightarrow$  large higher order corrections

$$m_{h_1}^2 \approx M_Z^2 \cos^2 2\beta \left( 1 - \frac{3}{8\pi^2} \frac{m_t^2}{v^2} \ln \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{M_t^2} \right) + \frac{3}{4\pi^2} \frac{m_t^4}{v^2} \ln \frac{m_{\tilde{t}_1} m_{\tilde{t}_2}}{M_t^2} + \dots$$

- ▶  $\Rightarrow$  also large corrections to prediction for  $M_Z$  at SUSY scale  $M_S$ :

$$\frac{M_Z^2}{2} = -\mu^2 + \overbrace{\frac{m_{H_d}^2 - m_{H_u}^2}{\tan^2 \beta - 1} \tan^2 \beta}^{\text{RGE effects}} + \delta_{1\text{-loop}},$$

$$\delta_{1\text{-loop}} = \frac{3}{8\pi^2} \frac{m_t^2}{v^2 \cos 2\beta} \left[ m_{\tilde{t}_1}^2 \left( \ln \frac{m_{\tilde{t}_1}^2}{M_S^2} - 1 \right) + m_{\tilde{t}_2}^2 \left( \ln \frac{m_{\tilde{t}_2}^2}{M_S^2} - 1 \right) \right] + \dots$$

- ▶  $\Rightarrow$  Little Hierarchy Problem
- ▶  $\mu$ -problem:  $M_Z^2 = -\mu^2 + \dots \Rightarrow \mu \sim$  soft parameters?

# Solving the Little Hierarchy Problem?

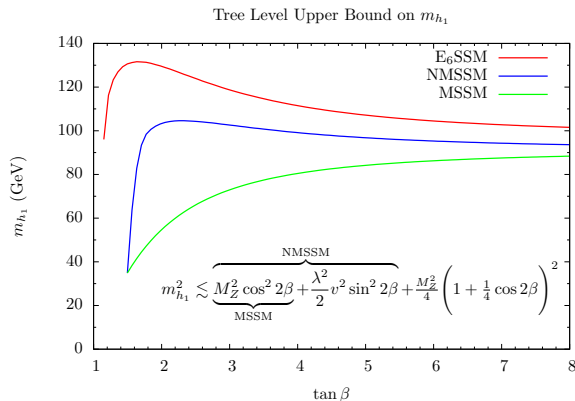
- Additional  $F$ -terms from new fields, e.g. NMSSM:

$$W \supset \lambda \hat{S} \hat{H}_d \cdot \hat{H}_u$$

- Additional  $D$ -terms from extended gauge sector, e.g.  $U(1)$  extensions:

$$\begin{aligned} & SU(3)_C \times SU(2)_L \times U(1)_Y \\ & \rightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)' \end{aligned}$$

- Care required (anomaly cancellation)



## $E_6$ Inspired Models

- ▶ Can arise from  $E_8 \times E'_8$  heterotic string theory
- ▶ Low-energy models have extra  $U(1)$  coming from breakdown

$$\begin{aligned} E_6 &\longrightarrow SO(10) \times U(1)_\psi \\ &\longrightarrow SU(5) \times U(1)_\psi \times U(1)_\chi \\ &\longrightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)_\psi \times U(1)_\chi \\ &\longrightarrow SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)' \end{aligned}$$

- ▶ Resulting  $U(1)' = U(1)_\chi \cos \theta_{E_6} + U(1)_\psi \sin \theta_{E_6}$
- ▶ Matter content fills complete **27** representations
  - ▶  $\Rightarrow$  anomaly cancellation guaranteed
  - ▶ Exotic states at low-energies (extra scalars + vector-like fermions, anyone?)
- ▶ Break  $U(1)'$  with singlet  $\Rightarrow$  dynamically generate  $\mu$  term, massive  $Z'$

# The E<sub>6</sub>SSM

- ▶ Additional  $U(1)_N$  under which right-handed neutrinos are uncharged ( $\theta_{E_6} = \arctan \sqrt{15}$ )
- ▶ Extra  $\hat{L}_4, \hat{\bar{L}}_4$  from incomplete  $\mathbf{27}', \overline{\mathbf{27}}'$  for gauge unification
- ▶ Low-energy matter content from  $\mathbf{27}$ -plet:

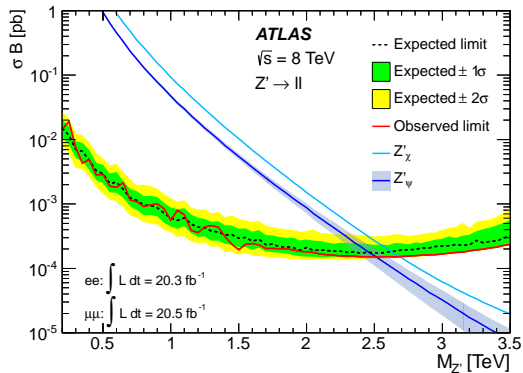
$$(\hat{Q}_i, \hat{u}_i^c, \hat{d}_i^c, \hat{L}_i, \hat{e}_i^c) + (\hat{D}_i, \hat{\bar{D}}_i) \\ + (\hat{S}_i) + (\hat{H}_i^u) + (\hat{H}_i^d)$$

- ▶ Higgs doublets  $\hat{H}_3^d \equiv \hat{H}_d, \hat{H}_3^u \equiv \hat{H}_u$  and one singlet  $\hat{S}_3 \equiv \hat{S}$  get VEVs ( $\Rightarrow$  EWSB and break  $U(1)_N$ )

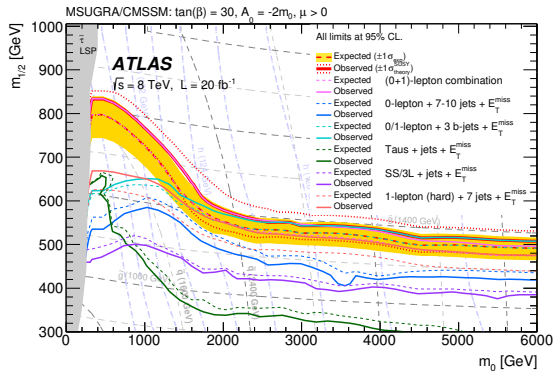
	$SU(3)_C$	$SU(2)_L$	$\sqrt{\frac{5}{3}} Q_i^Y$	$\sqrt{40} Q_i^N$
$\hat{Q}_i$	<b>3</b>	<b>2</b>	$\frac{1}{6}$	1
$\hat{u}_i^c$	$\overline{\mathbf{3}}$	<b>1</b>	$-\frac{2}{3}$	1
$\hat{d}_i^c$	$\overline{\mathbf{3}}$	<b>1</b>	$\frac{1}{3}$	2
$\hat{L}_i$	<b>1</b>	<b>2</b>	$-\frac{1}{2}$	2
$\hat{e}_i^c$	<b>1</b>	<b>1</b>	1	1
$\hat{S}_i$	<b>1</b>	<b>1</b>	0	5
$\hat{H}_i^u$	<b>1</b>	<b>2</b>	$\frac{1}{2}$	-2
$\hat{H}_i^d$	<b>1</b>	<b>2</b>	$-\frac{1}{2}$	-3
$\hat{D}$	<b>3</b>	<b>1</b>	$-\frac{1}{3}$	-2
$\hat{\bar{D}}$	$\overline{\mathbf{3}}$	<b>1</b>	$\frac{1}{3}$	-3
$\hat{L}_4$	<b>1</b>	<b>2</b>	$-\frac{1}{2}$	2
$\hat{\bar{L}}_4$	<b>1</b>	$\overline{\mathbf{2}}$	$\frac{1}{2}$	-2

$$W_{E_6SSM} \approx y_\tau \hat{L}_3 \cdot \hat{H}_d \hat{e}_3^c + y_b \hat{Q}_3 \cdot \hat{H}_d \hat{d}_3^c + y_t \hat{H}_u \cdot \hat{Q}_3 \hat{u}_3^c + \lambda_i \hat{S} \hat{H}_i^d \cdot \hat{H}_i^u + \kappa_i \hat{S} \hat{D}_i \hat{\bar{D}}_i + \mu' \hat{L}_4 \cdot \hat{\bar{L}}_4$$

# E<sub>6</sub>SSM: $M_{Z'} \sim M_S$



[PRD 90 (2014) 052005 (arXiv:1405.4123)]

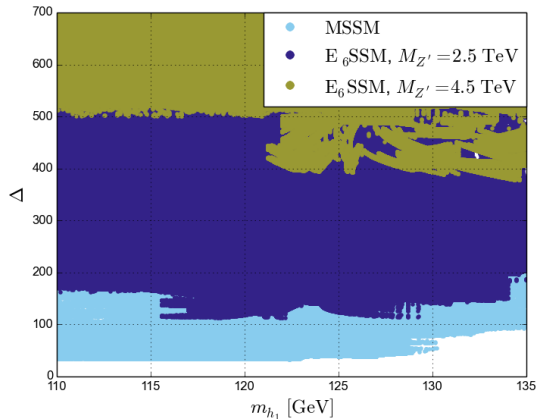


[JHEP 10 (2015) 054 (arXiv:1507.05525)]

# Issues in the E<sub>6</sub>SSM

$$\underbrace{c(\tan\beta)}_{O(1)} \frac{M_Z^2}{2} \approx -\lambda^2 \langle S \rangle^2 + \frac{m_{H_d}^2 - m_{H_u}^2 \tan^2 \beta}{\tan^2 \beta - 1} + d(\tan\beta) \frac{M_{Z'}^2}{2}$$

- ▶ Extra  $D$ -terms cut both ways
- ▶  $M_{Z'} \gtrsim 2.5$  TeV  $\Rightarrow$  another large contribution must cancel
- ▶ Unnatural or fine-tuned? E.g. as measured using traditional measure of fine-tuning
- ▶ Also, dangerous  $B$ ,  $L$  violating interactions  $\Rightarrow$  need approximate  $Z_2^H$  + exact  $Z_2^{L/B}$
- ▶ DM in simplest versions of model requires *another* exact  $Z_2^S$



# The SE<sub>6</sub>SSM

- ▶ Modification of E<sub>6</sub>SSM, arising from 5D or 6D orbifold GUT [1]
- ▶ Low-energy matter content augmented by components of extra **27'**,  $\overline{\mathbf{27}}'$ -plets, including:
  - ▶ doublets  $\hat{H}_u, \hat{H}_d$  (note: now not from **27**-plet)
  - ▶ singlets  $\hat{S}, \hat{\bar{S}}$  with opposite  $U(1)_N$  charges
- ▶ Still potentially dangerous  $B, L$  violating interactions
- ▶ But now can forbid using single, exact discrete  $\tilde{Z}_2^H$
- ▶ Stabilise scalar potential  $\Rightarrow$  pure singlet  $\hat{\phi}$

$$\begin{aligned}
 W_{\text{SE}_6\text{SSM}} = & \lambda \hat{S}(\hat{H}_d \cdot \hat{H}_u) - \sigma \hat{\phi} \hat{S} \hat{\bar{S}} + \frac{\kappa}{3} \hat{\phi}^3 + \frac{\mu}{2} \hat{\phi}^2 \\
 & + \Lambda_F \hat{\phi} + \lambda_{\alpha\beta} \hat{S}(\hat{H}_\alpha^d \cdot \hat{H}_\beta^u) + \kappa_{ij} \hat{S} \hat{D}_i \hat{\bar{D}}_j \\
 & + \tilde{f}_{i\alpha} \hat{S}_i(\hat{H}_u \cdot \hat{H}_\alpha^d) + f_{i\alpha} \hat{S}_i(\hat{H}_\alpha^u \cdot \hat{H}_d) \\
 & + g_{ij}^D(\hat{Q}_i \cdot \hat{L}_4) \hat{\bar{D}}_j + h_{i\alpha}^E \hat{e}_i^c(\hat{H}_\alpha^d \cdot \hat{L}_4) \\
 & + \mu_L(\hat{L}_4 \cdot \hat{\bar{L}}_4) + \tilde{\sigma} \hat{\phi}(\hat{L}_4 \cdot \hat{\bar{L}}_4) \\
 & + W_{\text{MSSM}}(\mu = 0)
 \end{aligned}$$

[1] R. Nevzorov, Phys. Rev. D 87, 015029 (2013) (arXiv:1205.5967)



# Dark Matter Candidates

	$\hat{Q}_i$	$\hat{u}_i^c$	$\hat{d}_i^c$	$\hat{L}_i$	$\hat{e}_i^c$	$\hat{N}_i^c$	$\hat{S}$	$\hat{\bar{S}}$	$\hat{S}_i$	$\hat{H}_u$	$\hat{H}_d$	$\hat{H}_\alpha^u$	$\hat{H}_\alpha^d$	$\hat{D}_i$	$\hat{\bar{D}}$	$\hat{L}_4$	$\hat{\bar{L}}_4$
$\tilde{Z}_2^H$	-	-	-	-	-	-	+	+	-	+	+	-	-	-	-	+	+
$Z_2^M$	-	-	-	-	-	-	+	+	+	+	+	+	+	+	+	-	-
$Z_2^E$	+	+	+	+	+	+	+	+	-	+	+	-	-	-	-	-	-

- Symmetry breaking pattern  $\Rightarrow$  automatically conserved  $Z_2^M = (-1)^{3(B-L)}$  ( $\equiv$  R-parity)
- Per MSSM, implies stable DM candidate (typically  $\tilde{\chi}_1^0$ )
- Write  $\tilde{Z}_2^H = Z_2^M \times Z_2^E \Rightarrow$  exotic  $Z_2^E$  also conserved
- Lightest exotic ( $Z_2^E$  odd) state is also stable DM candidate
- Very light inert singlinos form hot DM, but negligible contribution to relic density
- $\therefore$  relic density almost entirely due to ordinary neutralino

## EWSB in the $SE_6$ SSM

- ▶  $E_6$ SSM: EWSB conditions  $\Rightarrow M_{Z'} \sim M_S$  (i.e.  $\langle S \rangle \sim M_S$ )
- ▶ At physical minimum,

$$\langle H_d^0 \rangle = \frac{v_1}{\sqrt{2}}, \quad \langle H_u^0 \rangle = \frac{v_2}{\sqrt{2}}, \quad \langle S \rangle = \frac{s_1}{\sqrt{2}}, \quad \langle \bar{S} \rangle = \frac{s_s}{\sqrt{2}}, \quad \langle \phi \rangle = \frac{\varphi}{\sqrt{2}}$$

- ▶  $S, \bar{S}$  VEVs develop along nearly  $D$ -flat direction  $\langle S \rangle = \langle \bar{S} \rangle$  with

$$\langle S \rangle \approx \langle \bar{S} \rangle \sim \frac{M_S}{\sigma}$$

- ▶ Small  $\sigma \Rightarrow M_{Z'}^2 \sim g_1'^2 Q_S^2 (s_1^2 + s_2^2)$  far heavier than  $M_S$
- ▶  $\therefore$  can have  $M_{Z'}$  far above limits while keeping sparticles (relatively) light
- ▶ Note: also suppress  $D$ -term contribution to EWSB scale, i.e.  $M_{Z'}^2$

# The CSE<sub>6</sub>SSM

- ▶ General model is complicated
  - ▶  $O(200)$  new parameters (assuming no new sources of CP-violation)
  - ▶ Many masses and mixings
- ▶ Consider constrained model (CSE<sub>6</sub>SSM) inspired by gravity mediated SUSY breaking
- ▶ Universal soft breaking parameters:  $M_{1/2}$ ,  $A_0$ ,  $B_0$ ,  $m_0$
- ▶ Want  $Z'$  mass decoupled from EWSB conditions  $\Rightarrow$  fix heavy  $s = \sqrt{s_1^2 + s_2^2} = 650$  TeV
- ▶ Suitable DM candidate  $\Rightarrow$  small  $\mu_{\text{eff}}$ , i.e. small  $\lambda$
- ▶ Exotic  $\lambda_{\alpha\beta}$ ,  $\kappa_i$ , ... free, but small  $\lambda \Rightarrow$  natural for these to be small also
- ▶ FlexibleSUSY  $\Rightarrow$  mass spectrum at full one-loop level, including full two-loop RGEs
  - ▶ In general, want to use automated tools for studying with high precision

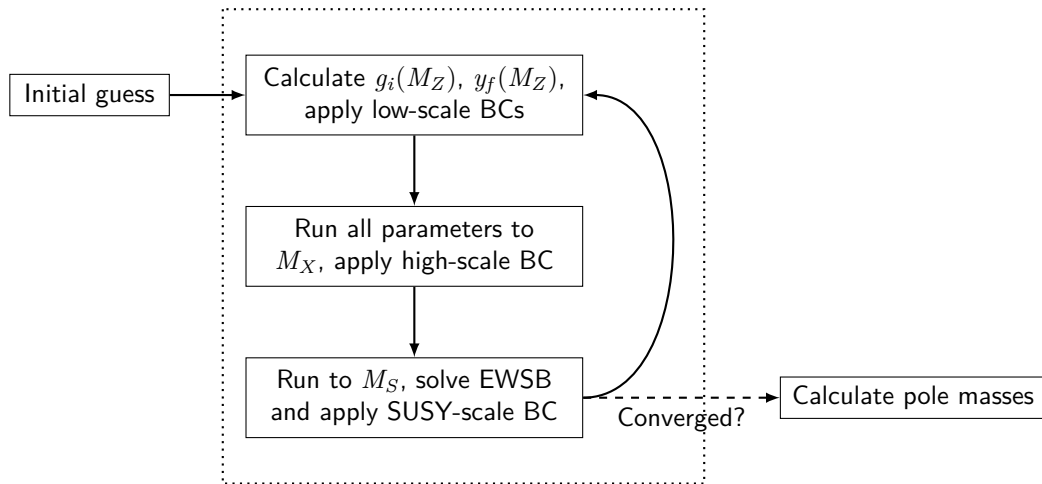
# Semi-analytic Solutions

- ▶ Useful technique for studying constrained models (e.g. CE<sub>6</sub>SSM)
- ▶ RGEs for SUSY parameters independent of soft SUSY breaking parameters (up to effects of threshold corrections)
- ▶ Treat RGEs for SUSY breaking parameters as system of linear ODEs
- ▶ Combine with boundary conditions  $\Rightarrow$  semi-analytic solutions,

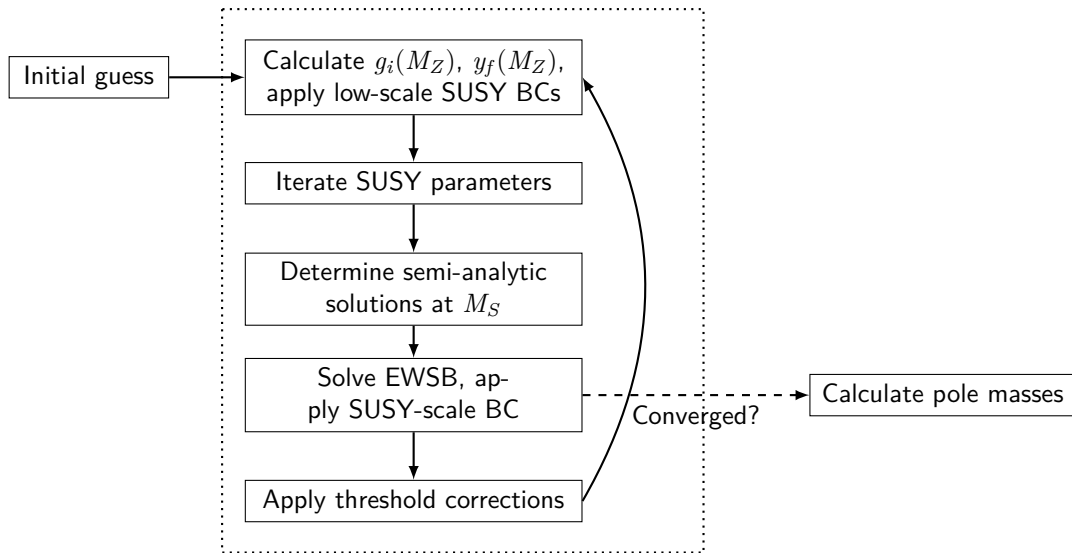
$$\begin{aligned} M_i(Q) &= p_i(Q)M_{1/2} + q_i(Q)A_0, & A_i(Q) &= e_i(Q)A_0 + f_i(Q)M_{1/2}, \\ m_i^2(Q) &= a_i(Q)m_0^2 + b_i(Q)M_{1/2}^2 + c_i(Q)A_0M_{1/2} + d_i(Q)A_0^2, \dots \end{aligned}$$

- ▶ Coefficients depend only on SUSY parameters, calculated numerically
- ▶ Relate SUSY scale parameters to high-scale inputs, e.g. trade  $m_0$  for  $\mu_{\text{eff}}$ 
  - ▶ Contrast with e.g. CMSSM:  $\mu$  output for given  $m_0$

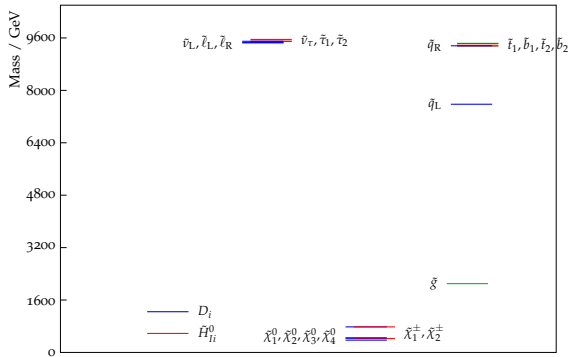
## Two-scale Algorithm



# Semi-analytic Algorithm

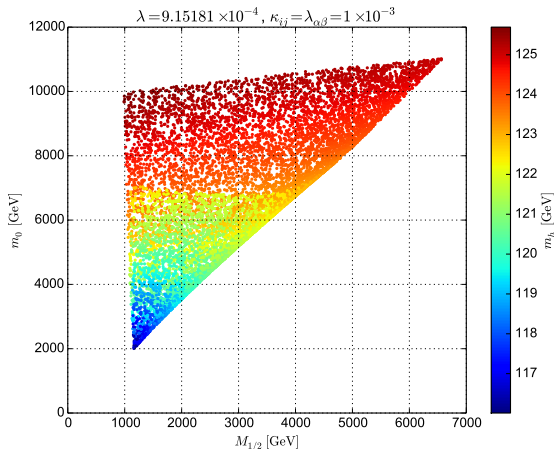


# Sparticle Masses



- ▶ Altered RG running  $\Rightarrow$  gaugino masses at low energies lighter than sfermion masses
- ▶ EWSB conditions  $\Rightarrow$  physical solutions satisfy  $m_0 > M_{1/2}, A_0$
- ▶  $\therefore$  MSSM sfermions are very heavy, out of reach of LHC run II
- ▶ Conversely, small  $\lambda_{\alpha\beta}, \kappa_i \Rightarrow$  light exotic fermions are observable
  - ▶ Exotic leptoquarks  $D_i$ , e.g. in  $pp \rightarrow t \bar{t} \tau^+ \tau^- + E_T^{miss} + X$ ,  $pp \rightarrow b \bar{b} + E_T^{miss} + X$
  - ▶ Charged and neutral inert Higgsinos  $pp \rightarrow Z Z + E_T^{miss} + X$ ,  $pp \rightarrow W Z + E_T^{miss} + X$ ,  $pp \rightarrow W W + E_T^{miss} + X$

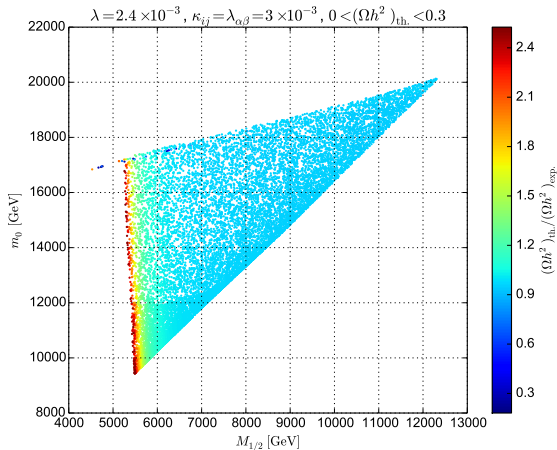
# CP-even Higgs Mass



- ▶ Split spectrum  $\Rightarrow$  large logarithms contribute to  $m_h$ 
  - ▶ Small exotic couplings  $\Rightarrow$  corresponding logarithms small
  - ▶ Mainly due to  $\tilde{t}$ 's
- ▶ Standard  $\overline{\text{DR}}$  calculation is inaccurate
- ▶ Solution: use EFT calculation (using SUSYHD now, cross checked with FlexibleHiggs prototype)
- ▶  $m_h \sim 125$  GeV obtained over large part of parameter space



# Dark Matter Relic Density: Higgsino LSP

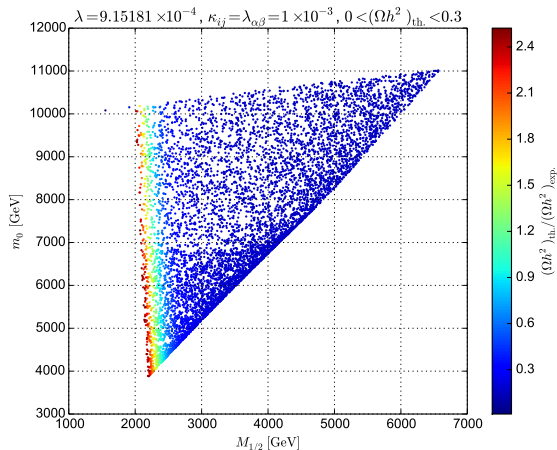


- For Higgsino LSP,

$$\Omega h^2 \sim (\Omega h^2)_{\text{obs}} \left( \frac{\mu_{\text{eff}}}{1 \text{ TeV}} \right)$$

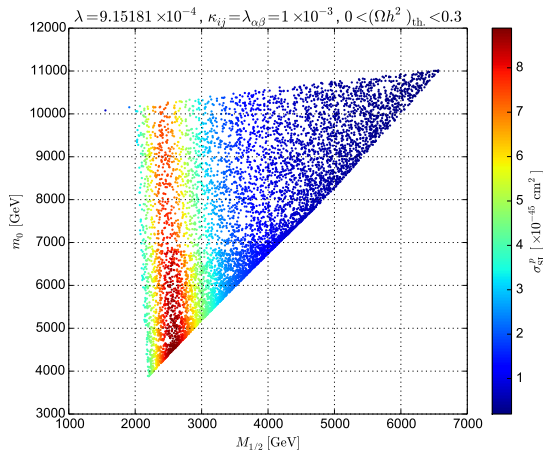
- Mainly co-annihilations  $\tilde{\chi}_i^0 \tilde{\chi}_1^\pm \rightarrow \bar{f}_1 f_2$ ,  
 $\tilde{\chi}_1^0 \tilde{\chi}_2^0 \rightarrow f\bar{f}$
- Requires  $m_{\tilde{\chi}_1^0} \sim 1 \text{ TeV}$  Higgsino
- $M_{1/2}$  must be large ( $\gtrsim 6 \text{ TeV}$ ) to suppress bino component
- Universal gaugino masses at  $M_{\text{GUT}} \Rightarrow$  all gauginos are heavy, e.g.  $m_{\tilde{g}} \gtrsim 5 \text{ TeV}$

# Dark Matter Relic Density: Mixed Bino-Higgsino LSP



- ▶ Alternative: mixed bino-Higgsino LSP, e.g.  $m_{\tilde{\chi}_1^0} \sim 300$  GeV
- ▶ Need  $M_1 \sim \mu_{\text{eff}}$  for large mixing
  - ▶  $M_1 < \mu_{\text{eff}} \Rightarrow$  pure bino, Universe overclosed
  - ▶  $M_1 > \mu_{\text{eff}} \Rightarrow \tilde{\chi}_1^0$  light Higgsino LSP, underabundant DM
- ▶ Mainly pair annihilation, e.g.  $\tilde{\chi}_1^0 \tilde{\chi}_1^0 \rightarrow \bar{f} f$
- ▶ Much lighter  $M_{1/2} \Rightarrow$  can also have  $\tilde{g}$  within reach at run II and HL-LHC

# Direct Detection



- ▶ Strong constraints from LUX limits on  $\sigma_{SI}$
- ▶  $\sigma_{SI} \propto$  (product of Higgsino, bino components)
- ▶  $\therefore$  highly mixed bino-Higgsino LSP  $\Rightarrow$  large  $\sigma_{SI}$
- ▶ Very close to LUX limit, would be immediately observed at XENON1T
- ▶ Pure Higgsino LSP  $\Rightarrow$  small bino component (+ heavy),  $\sigma_{SI}$  is safely below LUX limits
- ▶ But still detectable at XENON1T

# Summary

- ▶  $E_6$  inspired models can resolve the Little Hierarchy problem of MSSM
- ▶ Simplest variants, e.g.  $E_6$ SSM, faced with several issues ( $Z'$  limits, multiple discrete symmetries ...)
- ▶ Considered well-motivated extension,  $SE_6$ SSM, with exact custodial symmetry
- ▶ Heavy  $Z' \Rightarrow$  sfermions out of LHC reach, but observable exotic fermions, gauginos
- ▶ DM relic density can be fitted by MSSM-like neutralino
  - ▶ Higgsino LSP ( $\Rightarrow$  heavy spectrum)
  - ▶ Mixed bino-Higgsino LSP ( $\Rightarrow$  observable gauginos)
- ▶ Direct detection limits  $\Rightarrow$  strong constraints, easily observable at XENON1T

Thank you for listening!